

# Space-time modelling and simulation of extreme rainfall

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with the members of the Cerise and Fraise lefe projects !

MASCOT 2021 Meeting



# Proposition of a hierarchical space-time model for extreme precipitation data



$s = \text{"Montpellier"} , t = \text{"30 Sept. 2014"}$



$s = \text{"Venice"} , t = \text{"8 Nov. 2014"}$

Avignon



T Opitz

Montpellier



JN Bacro

Venice



C Gaetan

Cerise and Fraise LEFE projects

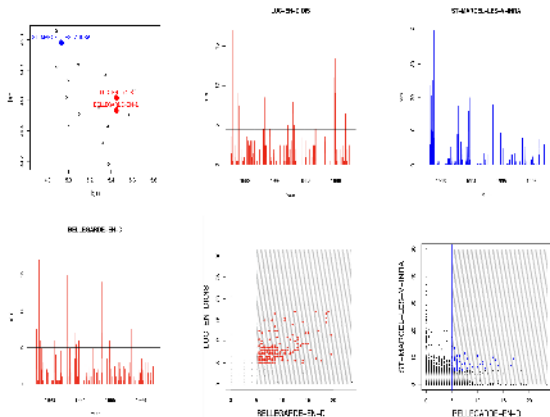


# Rainfall data

- ▶ Hourly observations at 50 rainfall stations for the years 1993 to 2014 from September to November (54542 hours)
- ▶ Moderately large dataset (50 × 54542 observations)



# Three stations in France



- ▶ clusters of strong values over space and time,
- ▶ strong variations at very small spatial and temporal scales

## Bivariate max-stable distributions

Let  $(X_i, Y_i) \sim F$  be independent random vectors with w.l.g. unit Fréchet margins  $K(x) = \exp(-1/x)$ ,  $x > 0$ . If a non-degenerate limit distribution for  $(M_{x,n}, M_{y,n}) = (\max_{i=1, \dots, n} X_i, \max_{i=1, \dots, n} Y_i)$  exists ( $F \in D(G)$ ),

$$\lim_{n \rightarrow \infty} \mathbb{P}(M_{x,n} \leq nx, M_{y,n} \leq ny) = G(x, y)$$

then  $G$  is **max-stable**:  $G^k(kx_1, ky_2) = G(x, y)$

- If

$$G(x, y) = K(x)K(y) = \exp\left(-\frac{1}{x}\right)\exp\left(-\frac{1}{y}\right)$$

$\hookrightarrow$  *ultimately, normalized maxima of  $X$  and  $Y$  are independent.*

$(X, Y)$  are said to be **Asymptotically Independent (AI)**.

Otherwise,  $(X, Y)$  are **Asymptotically Dependent (AD)**.

## Dependence measures $\chi$ and $\bar{\chi}$

Let  $(X, Y) \sim F \in D(G)$ , with  $F_X$  and  $F_Y$  margins.

### The $\chi$ parameter

$$\begin{aligned}\chi &= \lim_{u \rightarrow 1} \mathbb{P}(F_Y(Y) > u | F_X(X) > u) \\ &= \lim_{u \rightarrow 1} 2 - \frac{\log \mathbb{P}(F_X(X) \leq u, F_Y(Y) \leq u)}{\log \mathbb{P}(F_X(X) \leq u)} \\ &\equiv \lim_{u \rightarrow 1} \chi(u)\end{aligned}$$

- $\chi > 0 \Rightarrow X$  and  $Y$  are **AD**;  
 $\hookrightarrow \chi$  quantifies the strength of the extremal dependence.
- $\chi = 0 \Rightarrow X$  and  $Y$  are **AI**.  
 $\hookrightarrow \chi$  unable to provide dependence information for AI case !

### The $\bar{\chi}$ parameter

$$\begin{aligned}\bar{\chi} &= \lim_{u \rightarrow 1} \frac{2 \log \mathbb{P}(F_X(X) > u)}{\log \mathbb{P}(F_X(X) > u, F_Y(Y) > u)} - 1 \\ &\equiv \lim_{u \rightarrow 1} \bar{\chi}(u)\end{aligned}$$

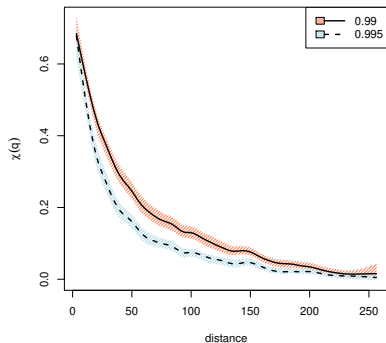
- $\bar{\chi} = 1 \Rightarrow X$  and  $Y$  are **AD**.
- $-1 \leq \bar{\chi} < 1 \Rightarrow X$  and  $Y$  are **AI**;  
moreover  $\bar{\chi}$  provides a measure that increases with dependence strength.

**Example 1:** Gaussian vectors with correlation parameter  $\rho \neq 1$ :  $\chi = 0$ ,  $\bar{\chi} = \rho$ .

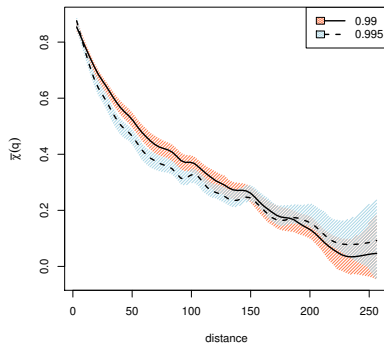
**Example 2:** For **max-stable distribution**,  $\chi(u) = \chi$  (same dependence structure  $\forall u$ )

# Our rainfall data: extremal dependence measure I

Spatial lag:  $x = (s, t), x' = (s + h_s, t)$



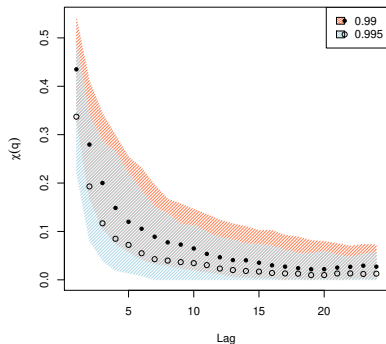
$\chi_{x,x'}(u)$



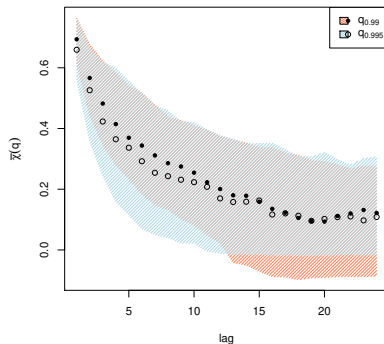
$\bar{\chi}_{x,x'}(u)$

# Our rainfall data: extremal dependence measure II

Temporal lag:  $x = (s, t), x' = (s, t + h_t)$



$\chi_{x, x'}(u)$



$\bar{\chi}_{x, x'}(u)$



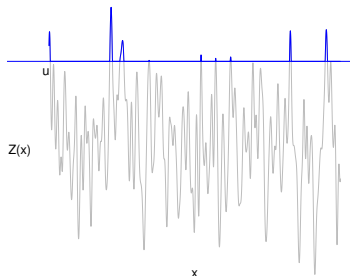
# Space-time setup

Bacro, J.N., Gaetan, C., Opitz, T., Toulemonde, G. (2020). "Hierarchical space-time modeling of exceedances with an application to rainfall data", *JASA*.

- ▶  $\{Z(x), x \in D\}$ , **space-time** process where  $x = (s, t)$ ,  $D \subset \mathbb{R}^2 \times \mathbb{R}_+$ 
  - ▶  $s$  space coordinate
  - ▶  $t$  time coordinate
- ▶ Types of concern when dealing with extreme values of the processes:
  - ▶ accurate inference for marginal distributions
  - ▶ assessment of the space-time dependence of the extreme values  
Possibly **asymptotically independent**
- ▶ What *Extreme* means for a process? no unique definition

# Exceedances

- ▶ Model for tail behaviour of  $Z(x)$  by fixing a “high” threshold  $u$  and focusing only on the (left-censored) values above  $u$  (**exceedances**)



↪ We explicitly model the original event data

## Marginal modelling: Generalized Pareto (GP) distribution

- ▶ Distribution for (censored) exceedances : the GP distribution
- ↪ asymptotic justification for  $u \rightarrow \tau_F$  (upper endpoint)

$$\begin{aligned}\Pr(Z(x) - u \leq y | Z(x) > u) &\simeq 1 - \left(1 + \xi \frac{y}{\psi}\right)_+^{-1/\xi} \\ &:= H(y; \xi, \psi), \quad y \geq 0\end{aligned}$$

- ▶ A different look at the GP distribution (when  $\xi > 0$ ): GP distribution **as a Gamma mixture for the rate of the exponential distribution**:

$$V | G \sim \text{Exp}(G), \quad G \sim \text{Gamma}(1/\xi, \psi/\xi) \Rightarrow V \sim \text{GP}(\cdot; \xi; \psi).$$

## Hierarchical space-time process with GP marginals

**Hierarchical** formulation for exceedances (following an idea of Bortot and Gaetan, 2014)

$$Y(x) := (Z(x) - u) \mathbb{1}\{Z(x) > u\}$$

- ▶ latent space-time process *with Gamma marginals*

$$\mathbf{G}(x) \sim \text{Gamma}(\alpha, \beta)$$

$$\hookrightarrow Y(x) | [\mathbf{G}(x), Y(x) > 0] \sim \text{Exp}(\mathbf{G}(x))$$

$$P(Y(x) > 0 | \mathbf{G}(x)) = \exp(-\kappa \mathbf{G}(x))$$

where  $\kappa > 0$  is a parameter controlling the rate of upcrossings of the threshold.

↔ joint space-time structure of the zero part and the positive part in the distribution of  $Y(x)$

## Multivariate distribution over the threshold

Exploiting a direct connection between probabilities for  $Y(\cdot)$  and  $\mathcal{L}_G(\cdot)$ , we obtain:

$$\Pr(Z(x) > \mathbf{u}) = \mathbb{E}[\Pr(Z(x) > \mathbf{u} \mid G(x))] = \mathcal{L}_{G(x)}(\boldsymbol{\kappa})$$

↪ Data  $\mathbf{z} = (z_1, \dots, z_n)'$  ; for  $\mathbf{z} \geq \mathbf{u}$ ,

$$\Pr(Z(x_1) > z_1, \dots, Z(x_n) > z_n) = \mathcal{L}_G(\mathbf{z} - (\mathbf{u} - \boldsymbol{\kappa}))$$

↪ Multivariate densities can be evaluated as soon as  $\mathcal{L}_G$  is known.

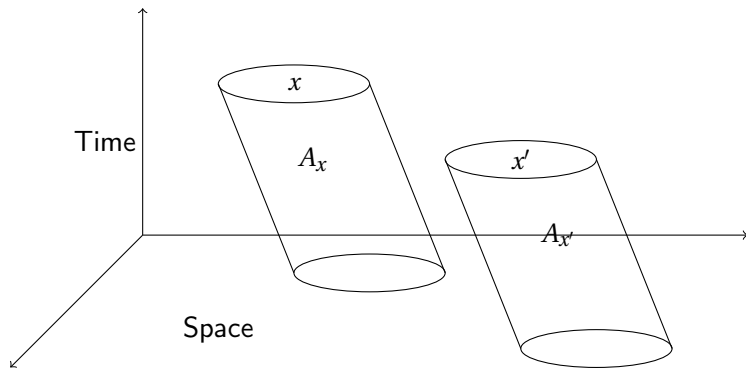
▶ Notation for **bivariate distribution** with  $z_1 > u$  and  $z_2 > u$  :  
 $\Pr(Z(x_i) \leq z_i, Z(x_j) \leq z_j) = H(z_i, z_j)$

# Which space-time process $G(\cdot)$ with Gamma marginals?

Based on **Slated elliptical cylinder** + **Gamma random field**

(Huser and Davison, 2014)

(Wolpert and Ickstadt, 1998)



- ▶ The **ellipse** describes the spatial influence zone of a storm
- ▶ The ellipse (storm) moves through space with a **velocity**  $\omega$
- ▶ The **duration** of a storm is  $\delta > 0$

## Which space-time process $G(\cdot)$ with Gamma marginals?

We propose to model the space-time process  $\{G(x), x \in \mathcal{D}\}$  as a **convolution** of a **Gamma random field**  $\Gamma(\cdot)$  (Wolpert and Ickstadt, 1998)

$$G(x) = \int_{A_x} \Gamma(dx') = \Gamma(A_x).$$

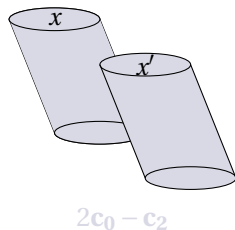
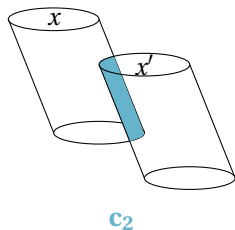
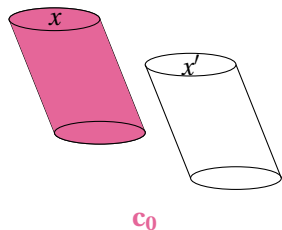
with  $\Gamma(\cdot)$  a Gamma RF defined on the space-time domain  $\mathcal{D} = \mathbb{R}^2 \times \mathbb{R}_+$

such that

- ★ for any set  $A$ ,  $\Gamma(A) := \int_A \Gamma(dx) \sim \text{Gamma}(\alpha(A), \beta)$ ;
- ★ for any sets  $A_1, A_2$  such that  $A_1 \cap A_2 = \emptyset$ ,  $\Gamma(A_1)$  and  $\Gamma(A_2)$  are **independent** random variables.

# Extremal dependence of $Z(\cdot)$ : Asymptotic Independence

$$\chi_{x,x'} = \mathbf{0} \text{ and } \bar{\chi}_{x,x'} = \frac{c_2}{2c_0 - c_2} \geq 0$$





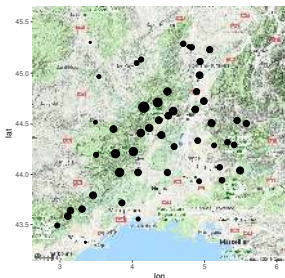
# Application to rainfall data

- ▶ Hourly observations at 50 rainfall stations for the years 1993 to 2014 from September to November (54542 hours)
- ▶ Moderately large dataset (50 × 54542 observations)

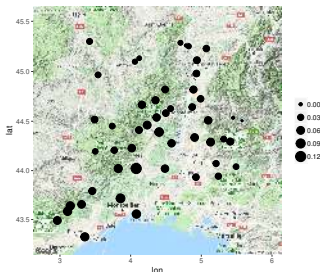


# Application to rainfall data

- ▶ Marginal distributions are not stationary in space



99% quantile

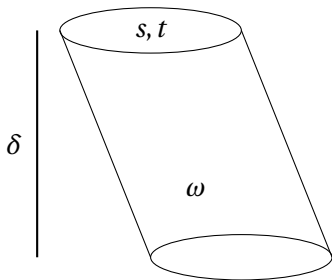
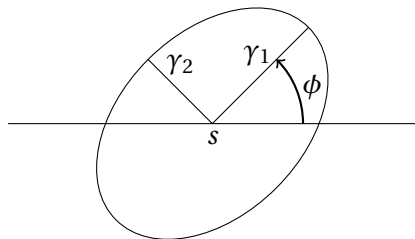


$\xi$  (tail parameter)

- ▶ Fit a GP distribution separately to each site  $s$  with thresholds chosen as the empirical quantiles  $q_{0.99}(s)$
- ▶ Transform the exceedances to the same GP distribution

## Space-time dependence parameters

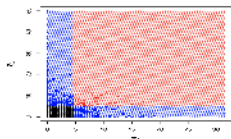
$$\theta = (\phi, \gamma_1, \gamma_2, \delta, \omega)'$$



## Inferential issues: composite likelihood approach

Let  $u$  be a sufficiently high threshold

- ▶ Different (censored) likelihood contribution  $L(z_1, z_2; \theta)$  of  $Z(x_1) = z_1, Z(x_2) = z_2$



- ▶ Weighted composite likelihood (Lindsay, 1988, Bevilacqua et al., 2012)

$$PL(\theta) = \prod_{i=1}^{n-1} \prod_{j=i+1}^n L(z_i, z_j; \theta) w_{ij}$$

$w_{ij}$  positive weights that depend on the space-time distance.

- ↪ Then we maximise pairwise weighted censored log-likelihood to obtain parameter estimations.

# Application to rainfall data

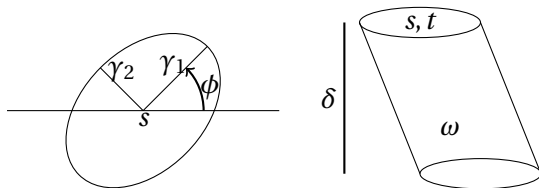
Two models for space-time dependence

**G1** Gamma-Pareto process

**G2** model G1 without velocity ( $\omega = 0$ )

Estimates, standard errors (in *italic*) values of the Gamma-Pareto fitted models.

$$\theta = (\phi, \gamma_1, \gamma_2, \delta, \omega)'$$



Model	Parameters					
	$\gamma_1$	$\gamma_2$	$\phi$	$\delta$	$\omega_1$	$\omega_2$
G1	165.062	318.823	0.085	20.184	0.723	0.446
	<i>23.459</i>	<i>19.811</i>	<i>0.026</i>	<i>0.948</i>	<i>0.195</i>	<i>0.009</i>
G2	175.817	294.323	0.041	20.036	0	0
	<i>11.879</i>	<i>25.291</i>	<i>0.064</i>	<i>1.039</i>	-	-

Parameter units are kms for  $\gamma_1$  and  $\gamma_2$ , radians for  $\phi$ , hours for  $\delta$  and kms per hour for  $\omega_1$  and  $\omega_2$ .

## Comparison with other AI processes

Comparison with three variants of a censored anisotropic Gaussian space-time copula.

**C1** Space-time separable model

**C2** Non-separable model (frozen field, Christakos, 2017)

**C3** Non-separable model with spherical correlation function

Comparison according to

▶ **CLIC** (minimum for our Gamma-Pareto process G1)

▶ **Bivariate conditional probabilities**

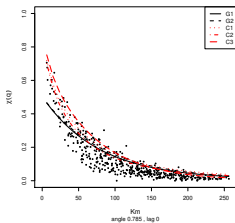
$$\Pr(Z(s, t) > q | Z(s', t - h_t) > q)$$

▶ **RMSE based on multivariate conditional probability**

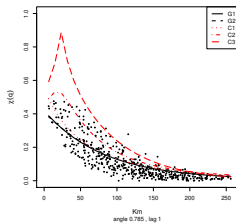
$$\chi_{s_i, h_t}^*(q) := \Pr(Z(s_j, t) > q, s_j \in \partial s_i | Z(s_i, t - h_t) > q)$$

# Angle $\pi/4$

lag 0



lag 1

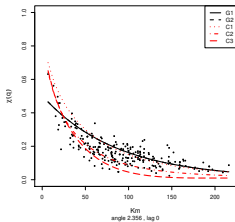


*Estimated probabilities  $\Pr(Z(s, t) > q | Z(s', t - h_t) > q)$  along different directions and at different temporal lags  $h_t$ . Dotted points correspond to empirical estimates. The value  $q$  is fixed to the empirical 99% quantile.*

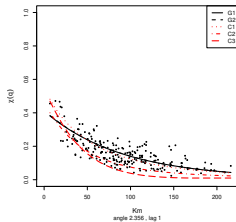


# Angle $3\pi/4$

lag 0



lag 1



*Estimated probabilities  $\Pr(Z(s, t) > q | Z(s', t - h_t) > q)$  along different directions and at different temporal lags  $h_t$ . Dotted points correspond to empirical estimates. The value  $q$  is fixed to the empirical 99% quantile.*

# RMSE

- ▶ Compute

- ▶ empirical estimates  $\hat{p}_i(h_t)$  of the multivariate conditional probability

$$\chi_{s_i; h_t}^*(q) := \Pr(Z(s_j, t) > q, s_j \in \partial s_i | Z(s_i, t - h_t) > q)$$

where  $\partial s_i$  is the set of the four nearest neighbors of site  $s_i$ ,  $i = 1, \dots, 50$ .

- ▶ Monte-Carlo estimates  $\tilde{p}_i^{(j)}(h_t)$ ,  $j = 1, \dots, 200$ .
- ▶ Calculate site-specific root mean squared errors (RMSE)

$$\text{RMSE}_i(h_t) = \left\{ \frac{\sum_{j=1}^{200} (\tilde{p}_i^{(j)}(h_t) - \hat{p}_i(h_t))^2}{200} \right\}^{1/2},$$

and the resulting total  $\text{RMSE}(\mathbf{h}_t) = \sum_{i=1}^{50} \text{RMSE}_i(\mathbf{h}_t)$ .

# RMSE

	RMSE(0)		RMSE(1)		RMSE(2)	
	$q_{0.99}$	$q_{0.995}$	$q_{0.99}$	$q_{0.995}$	$q_{0.99}$	$q_{0.995}$
G1	2.614	<b>2.096</b>	<b>1.901</b>	<b>1.643</b>	<b>1.475</b>	<b>1.496</b>
G2	2.605	<b>2.072</b>	<b>1.907</b>	<b>1.626</b>	<b>1.477</b>	<b>1.480</b>
C1	<b>2.240</b>	2.455	2.053	2.428	1.779	1.928

**Table 1:** Total root mean squared errors for the estimates of  $\chi_{s_i; h_t}^*(q)$  with  $h_t = 0, 1, 2$  hours.

## Conclusions on this part

- ▶ A **space-time model for threshold exceedances** of data with **asymptotically vanishing dependence strength** with **physical interpretation**.
- ▶ Extensions to asymptotic dependence are possible.
- ▶ Simulations of exceedances

## Why simulate extreme rainfall ?

Example of simulated water level and speed in Montpellier with a urban flood model.

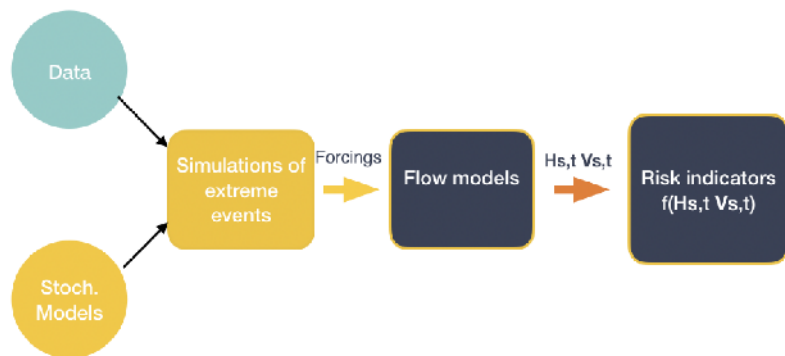


*Left : study area (600m x 600m). Center : simulated water depths. Right: detail view of the mesh. The lowest depths in blue and the largest depths (5 cm) in red*

- ↪ Input for **urban flood models**: **rainfall forcing**.
- ▶ Exploration of not already observed scenarios from limited observations
- ▶ Stochastic inputs for impact studies

**Reconstructing extreme space-time rainfall forcing scenarios as close to reality as possible is therefore a crucial issue.**

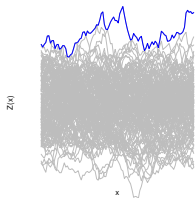
## Urban flood risk study



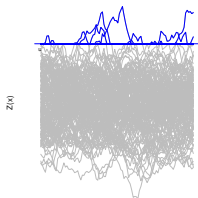
Toulemonde, G., Carreau, J., Guinot, V. (2020). "Space-time simulations of extreme rainfall : why and how ?" in S. Manou-Abi, S. Dabo-Niang, J. Salone (eds), Mathematical Modeling of Random and Deterministic Phenomena, Wiley.

# Which extremal behaviour of $Z = \{Z(x), x \in D\}$ ?

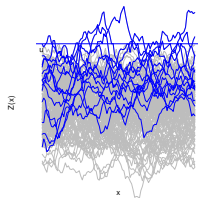
- ▶ what does it mean rainfall extreme we would like to simulate ?
- ▶ Events satisfying an exceedance condition



$\{\max_i Z_i(x)\}$   
Max-stable



$\{\max(Z(x), u)\}$   
Gamma-Pareto  
processes



$\{Z(x) | \sup_{x \in D} Z(x) > u\}$   
Pareto processes

$\ell$ -Pareto process  
 $\{Z(x), x \in D | \ell(Z(x)) > u\}$

## Semi-parametric simulation method

(Chailan, R., Toulemonde, G., and Bacro, J. (2017); Palacios-Rodriguez F., Toulemonde G., Carreau J., Opitz T. (2020))

### Construction of standard space-time $\ell$ -Pareto processes

(Based on Ferreira and de Haan, 2014; Dombry and Ribatet, 2015)

$$Z(s, t) := RY(s, t)$$

with  $R \sim \text{Pareto}(1, \gamma_R)$  independent of  $Y(s, t) \geq 0$ ,  $\ell(Y(s, t)) = 1$  with  $\ell$  a cost functional (a continuous non negative function that is homogeneous).



## Semi-parametric simulation method

- ▶ **Standardisation**  $\{Z^*(s, t), s \in \mathcal{S}, t \in \mathcal{T}\}$  the Pareto standardised process.
- ▶ **Extraction**
  - ▶ Defining extreme episodes  $\rightarrow$  Cost functional  $\ell +$  threshold  $u$
  - ▶ Select the  $m$  most extreme episodes  
 $\{Z_{[i]}^*(s, t), s \in \mathcal{S}_i \subset \mathcal{S}, t \in \mathcal{T}_i \subset \mathcal{T}\}, i \in \{1, \dots, m\}$

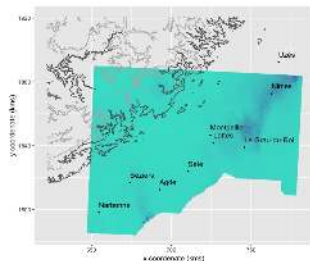
For each  $i \in \{1, \dots, m\}$ ,

- ▶ **Lifting procedure**
  - ▶ Non-parametric approach for the dependence structure
  - ▶ Sample  $R_i$  according to a Pareto r.v. with shape 1 and scale  $\alpha > 0$  and generate

$$V_i(s, t) = R_i \frac{Z_{[i]}^*(s, t)}{\ell(Z_{[i]}^*(s, t))} = R_i Y_i(s, t).$$

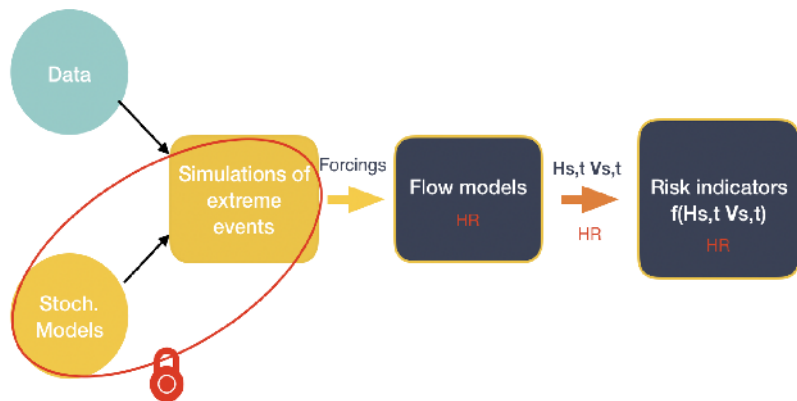
- ▶ **Back-transformation to original scale**

# Application to precipitation in Mediterranean France



- ▶ Reanalysis data-set
- ▶ Hourly rainfall totals ( $mm$ ).
- ▶  $133km \times 104km$  grid with  $1km$  spatial resolution.
- ▶ Years: from 1997 to 2007.  
 $N = 87642$  hours time steps.
- ▶  $\ell$ : Space-time neighborhoods( $15\text{ kms}, 24h$ )
- ▶  $u = 0.99$ -quantile.

## Some perspectives about urban flood risk study

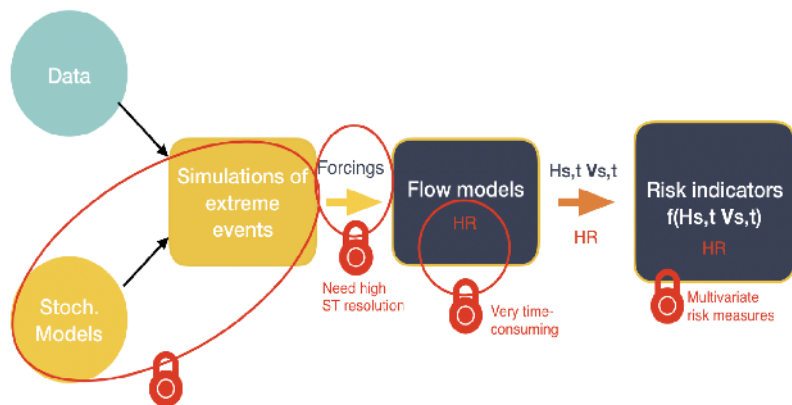


# Statistical modelling of extreme events

- ▶ Framework:
  - ▶ **multivariate,**
  - ▶ **temporal,**
  - ▶ **spatial**

→ taking into account associated complex dependence.
- ▶ Three main issues (I1), (I2) and (I3)
  - ▶ **(I1) Asymptotic independence** (hybrid according components)
  - ▶ **(I2) Spatial and/or temporal non-stationarity of the dependence structure** (Carreau J., Toulemonde G., 2020)
  - ▶ **(I3) Combination of extreme and non-extreme events.**

## Some perspectives about urban flood risk study



## Some references

- ▶ Bacro, J.N., Gaetan, C., Opitz, T., Toulemonde, G. (2020). Hierarchical space-time modeling of exceedances with an application to rainfall data, *JASA*, 115(530), 555-569.
- ▶ Carreau J., Toulemonde G. (2020). Spatial dependence structure for flood-risk rainfall. *Spatial Statistics*, 40.
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THANK YOU !